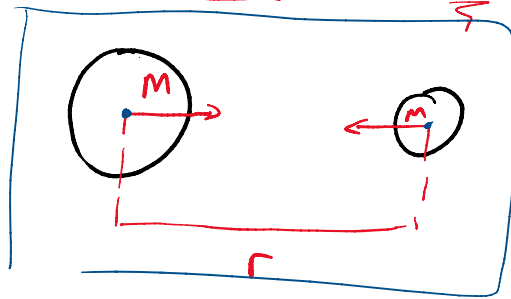


# Lecture 24

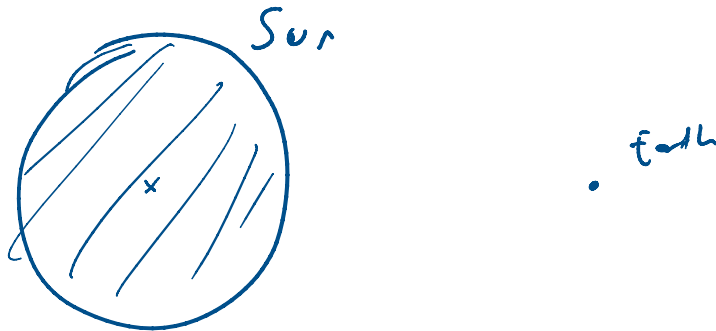
19 Aralık 2019 Perşembe 08:42

⇒ We Newtonian gravity  $\rightarrow \underline{|\vec{F}| = G \frac{Mm}{r^2}}$  ,  $U = -G \frac{Mm}{r}$

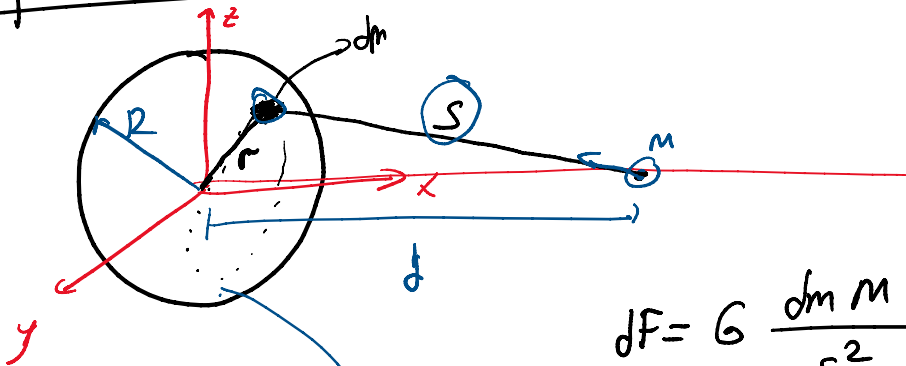


$$\nabla U = \vec{F}$$

- o Kepler's laws
- o Black holes, Einsteinian gravity

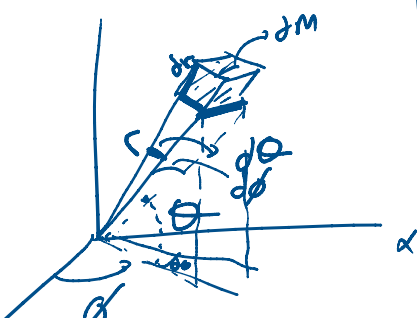


## Spherical Mass Distribution



$$dF = G \frac{dm M}{s^2}$$

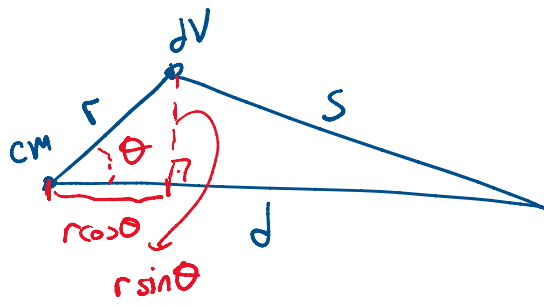
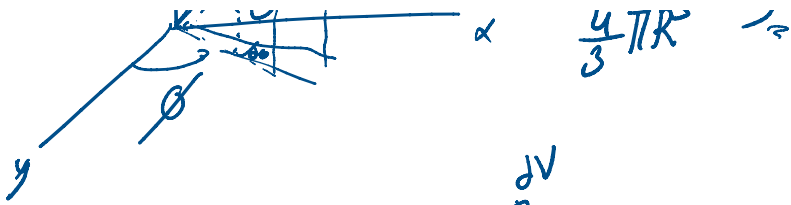
$$dU = -G \frac{dm M}{s}$$



$$\frac{M}{\frac{4}{3}\pi R^3} = \rho$$

$$dm = \rho \cdot dV = \rho \cdot \underbrace{dr r d\theta r \sin\theta \cdot d\phi}$$

0 . . . 1 . . . 12



$$s^2 = r^2 \sin^2 \theta + (d - r \cos \theta)^2$$

$$s^2 = r^2 \sin^2 \theta + d^2 - 2rd \cos \theta + r^2 \cos^2 \theta$$

$$s^2 = r^2 (\sin^2 \theta + \cos^2 \theta) - 2rd \cos \theta + d^2$$

Law's cosine  $\rightarrow$   $s^2 = r^2 - 2rd \cos \theta + d^2$

$$dU = -G \frac{m \rho r^2 \sin \theta dr d\theta d\phi}{(r^2 - 2rd \cos \theta + d^2)^{1/2}}$$

$$U = -G m \rho \int_0^R \int_{-\pi/2}^{\pi/2} \int_0^{2\pi} \frac{r^2 \sin \theta dr d\theta d\phi}{(r^2 - 2rd \cos \theta + d^2)^{1/2}}$$

$$\frac{d}{ds} f(s) = \frac{d\theta}{ds} \frac{d}{d\theta} f(s)$$

$$U = -\frac{G m \rho}{d} \iiint \frac{r ds dr d\theta}{s}$$

$$= -\frac{G m \rho 2\pi}{d} \int_0^R r dr \int_{-r+d}^{r+d} ds$$

$$= -\frac{G m \rho 2\pi}{d} \int_0^R \frac{r dr \cdot (r+d - (-r+d))}{2r^2 dr}$$

$$= \frac{G m \rho 2\pi}{d} \cdot R^3 = G m \rho \cdot \frac{4\pi}{3} R^3$$

$$\frac{d}{ds} s^2 = \frac{d}{d\theta} (r^2 - 2rd \cos \theta + d^2)$$

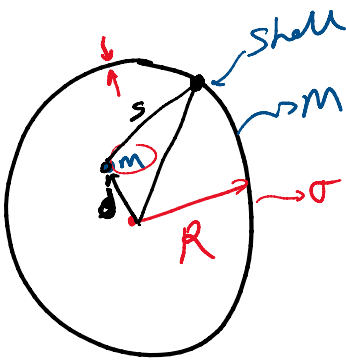
$$2s = \frac{d\theta}{ds} \cdot (2rd \sin \theta)$$

$$2s ds = 2rd \sin \theta d\theta \rightarrow \frac{s ds}{d} = r \sin \theta d\theta$$

$$= - \frac{G m \rho \cdot 2\pi \cdot \frac{4}{3} R^3}{d} = - \frac{G m \left( \rho \cdot \frac{4\pi}{3} R^3 \right)}{d}$$

$$= - G \frac{m M}{d}$$

⇒ Potential inside a spherical shell

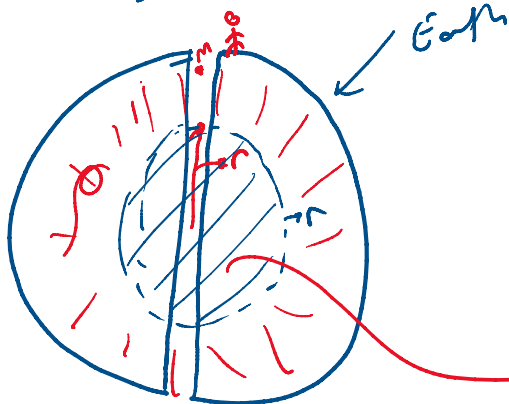


$$U = - \frac{G m \left( (\sigma \cdot dr) \cdot 2\pi R \right)}{2dR} \int_{R-d}^{R+d} ds$$

$$U = - G \frac{m M}{R}$$

$$\nabla U = 0$$

⇒ Journey through the center of the earth



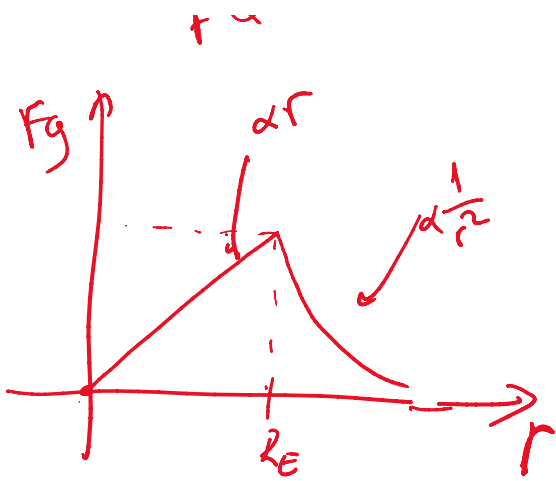
~ 12 km  
Russia

$$f_g = ?$$

$$m_r = \frac{V_r}{V_E} \cdot M_E \quad m_r = M_E \cdot \frac{r^3}{R_E^3}$$

$$f \propto x$$

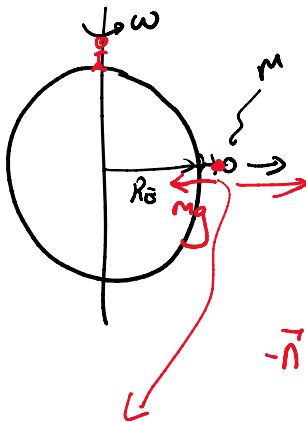
$$f_g = G \frac{M m}{r} = G \frac{M}{r} \cdot \left( \frac{M_E r^3}{\rho \cdot \frac{4}{3} \pi r^3} \right)$$



$$F_g = G \frac{Mm}{r^2} = G \frac{M}{r^2} \cdot \left( \frac{M_E r^3}{R_E^3} \right)$$

$$F_g = G \frac{M M_E \cdot r}{R_E^3}$$

⇒ Apparent weight & the earth's rotation



$$m \frac{v^2}{R_E} = m \omega^2 R_E$$

$$W_0 = mg$$

$$-N + mg = m \omega^2 R_E$$

$$W = mg - m \omega^2 R_E$$

apparent weight  $\sim 9.8 \text{ m/s}^2$

$0.0339 \text{ m/s}^2$

$$W \rightarrow g_e = 9.78 \dots$$

$$g_p = 9.81 \dots$$